

Propagation TeePee: A High Frequency (HF) Radio Spectral Feature Identified by Citizen Scientists

**Shing F. Fung¹, David Typinski², Richard Flagg³, Thomas Ashcraft⁴,
Wes Greenman⁵, Charles Higgins⁶, James Brown⁷, Larry Dodd⁸,
Francisco Reyes⁹, Jim Sky¹⁰, James Thieman¹¹ and Leonard Garcia¹²**

¹ITM Physics Laboratory, NASA Goddard Space Flight Center, Greenbelt MD 20771;

²AJ4CO Observatory, High Spring, FL 32655; ³RF Associates, Honolulu, HI 96826;

⁴Heliotown Observatory, Lamy, NM, 87540; ⁵LGM Radio Alachua, Alachua, FL 32615;

⁶Dept. Physics & Astronomy, Middle Tennessee State University, Murfreesboro, TN 37132 ;

⁷Hawks Nest Radio Astronomy Observatory, Industry, PA 15052;

⁸Georgia Amateur Radio Astronomy Observatory, Jasper, GA 30143;

⁹Department of Astronomy, U of Florida, Gainesville, FL 32611; ¹⁰Radio Sky Publishing, Louisville, KY 40214;

¹¹UMBC/NASA GSFC, Greenbelt MD 20771; ¹²SGT/NASA GSFC, Greenbelt MD 20771

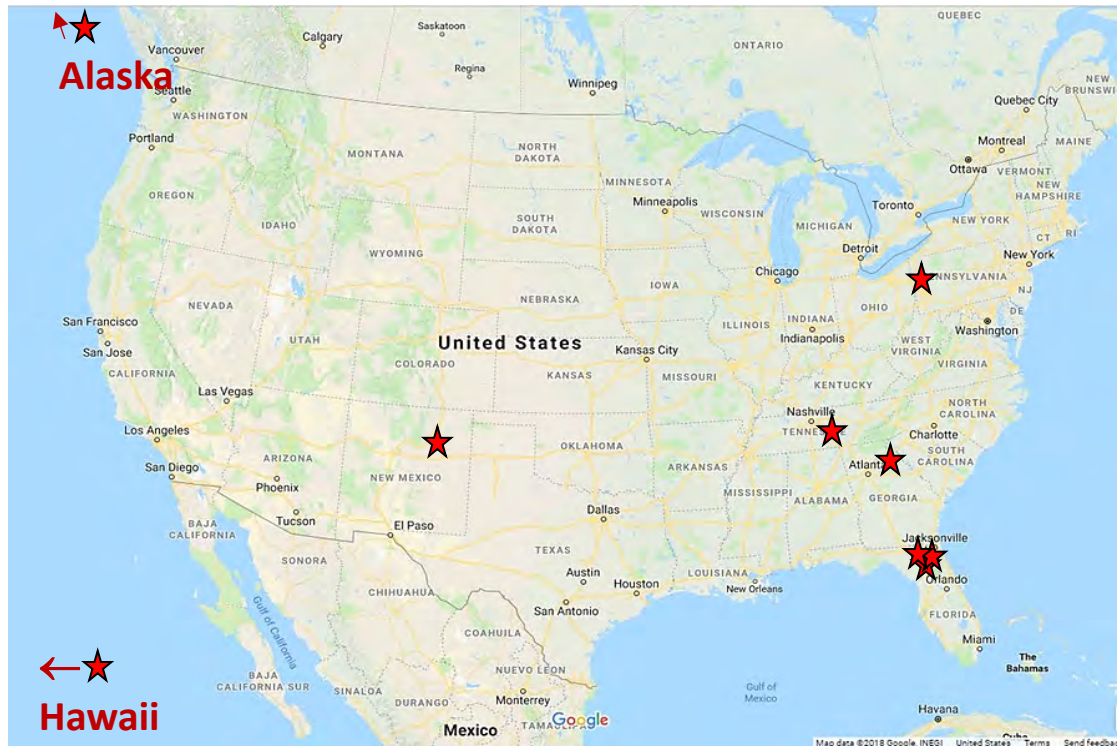


Citizen Science Under the Radio JOVE-NSSEC Partnership

- The Radio JOVE project <<http://radiojove.gsfc.nasa.gov/>>
 - Established for over 20 years as an informal education & outreach project
 - Over 2400 single-frequency kits distributed worldwide
 - Focuses on radio astronomy of Jupiter, the Sun, and the Galaxy
- The Spectrograph Users Group (SUG) is a more advanced subgroup that operates several radio spectrographs (15-30 MHz) across the continental US, Hawaii and Alaska.
- As a partner with the [NASA Space Science Education Consortium \(NSSEC\)](#) since 2016, the Radio JOVE project has been extended to perform citizen science research in heliophysics and space weather.

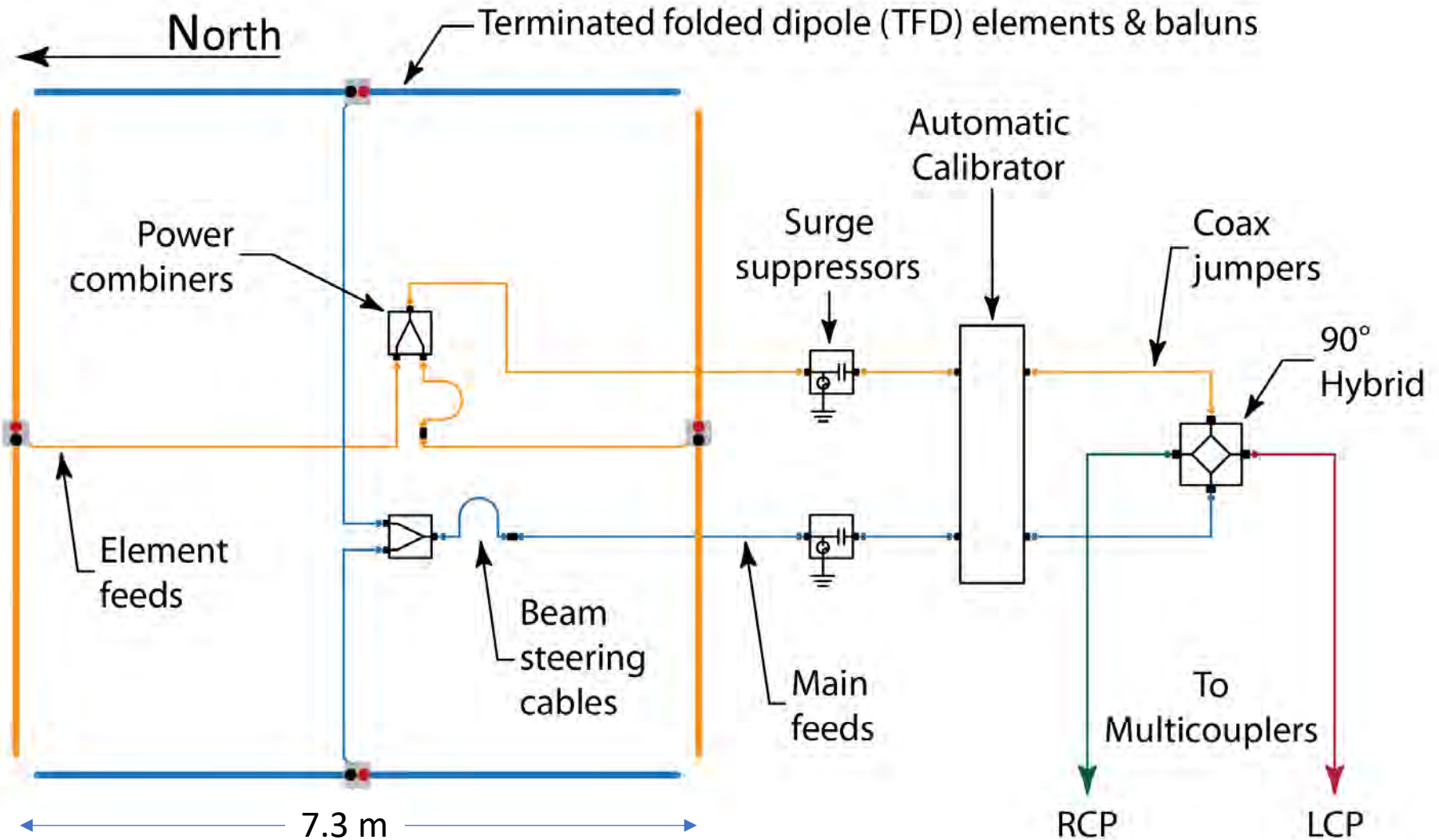
Radio JOVE Spectrograph Network

- Phase-1 network
 - 9 stations operating FSX spectrographs deployed across the Continental US, Alaska and Hawaii
- Phase-2 network
 - Under development
 - Based on SDRs [SDRPlay(2) & others]
 - Lower cost, more capable
 - More new stations to enable citizen science



Radio JOVE Spectrograph Users Group (SUG)

Typical Antenna (Phase 1 and 2)

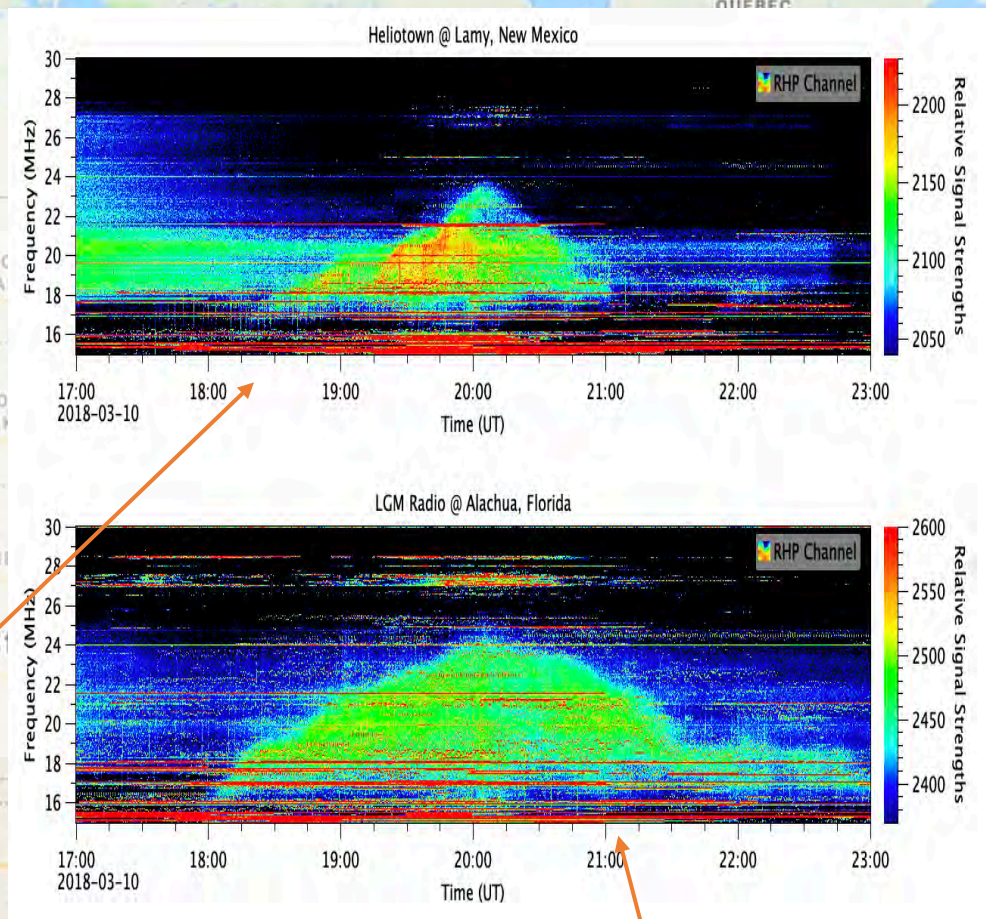


Radio JOVE Spectrograph Users Group (SUG) FSX Spectrograph

Frequency range:	15-30 MHz
Channel bandwidth:	30 kHz
Frequency step size:	50 kHz
Sweep rate:	2000 channels/s
Dwell time per channel:	0.5 ms
Number of channels:	300
Number of inputs:	2 (LCP & RCP)
Temporal resolution:	300 ms

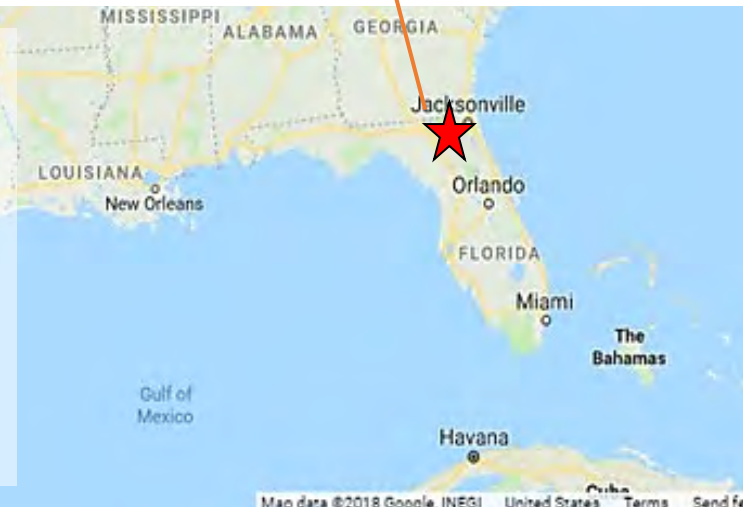
TeePee Spectral Characteristics:

Spectral enhancement with upper cutoff frequency first increases, then decreases with time, resulting in a spectral feature resembling a TeePee tent.

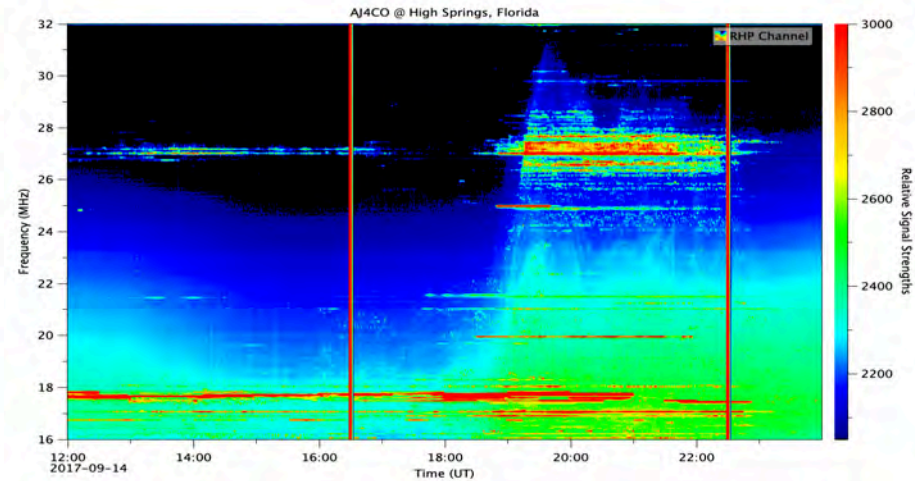
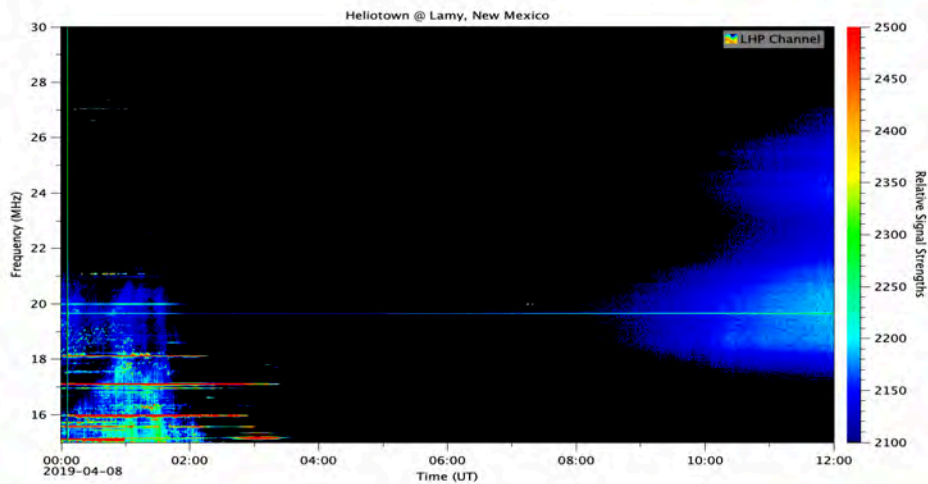


Examples show two TeePees observed simultaneously in Florida and New Mexico

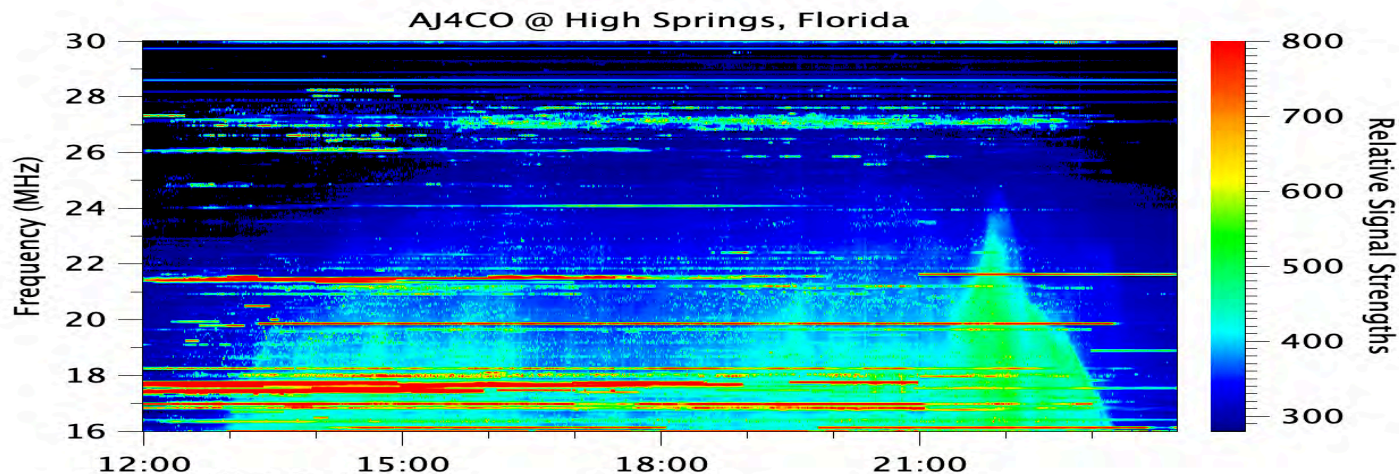
- Distance between the two observing stations ~2280 km, implying source range ~1000 km



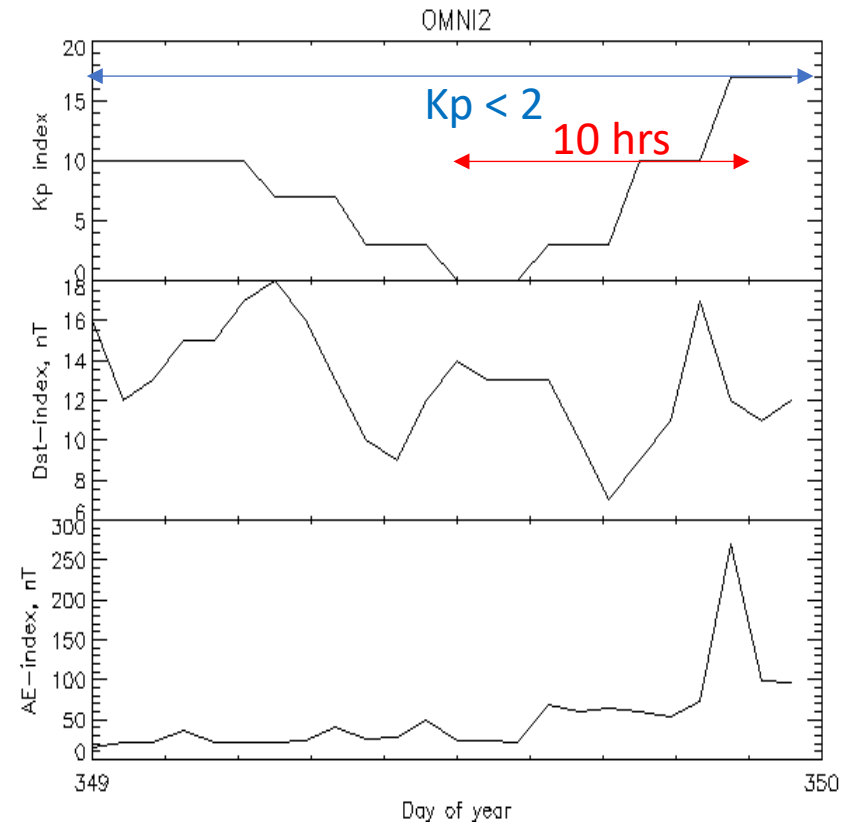
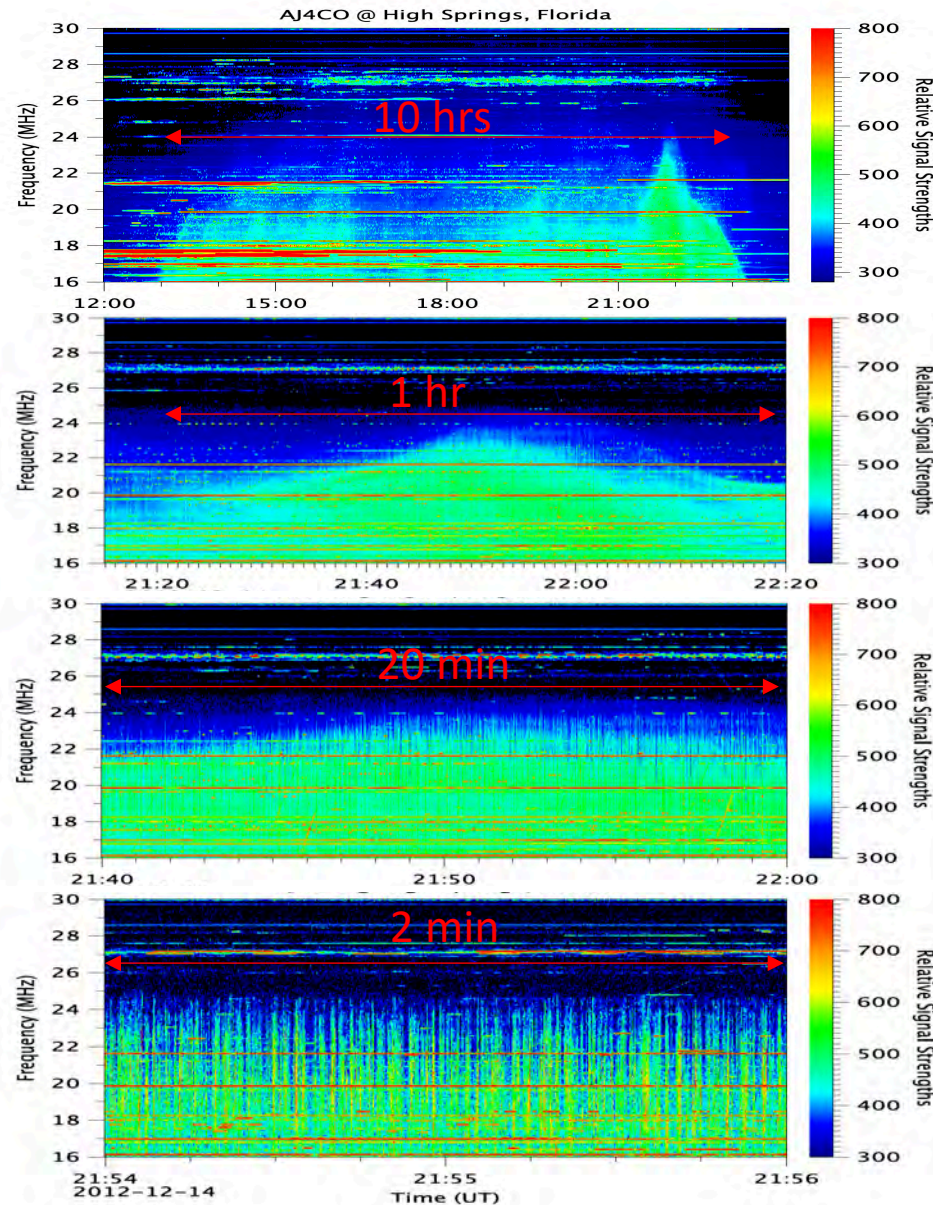
TeePees Often Appear in Groups



- Nested TeePees (above)
- TeePee Series (below)



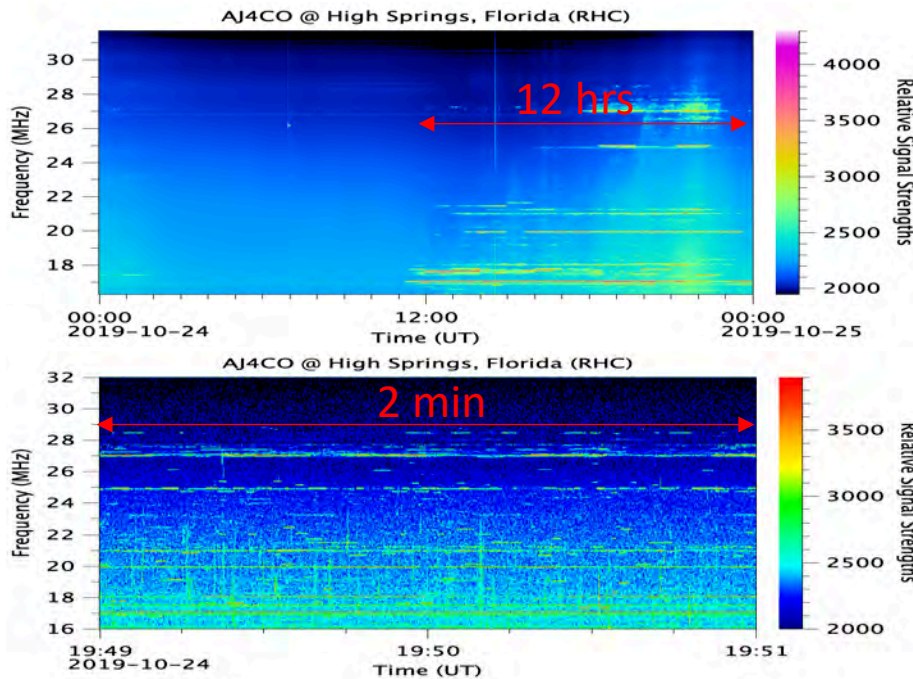
Zoomed-In View of Quiet-Time Event



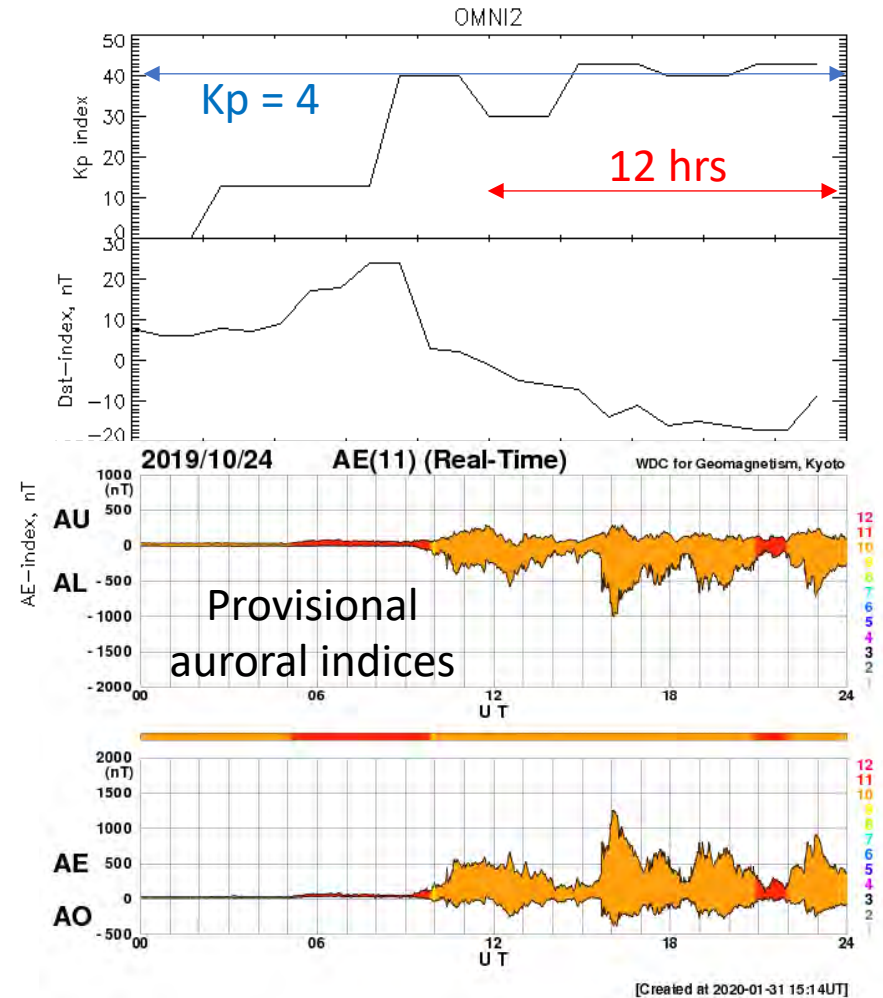
<https://omniweb.gsfc.nasa.gov/>

Discrete bursts seen occasionally during geomagnetically quiet times ($K_p < 2$) suggest lightning flashes might be the HF radiation source.

Event During “Disturbed Time”



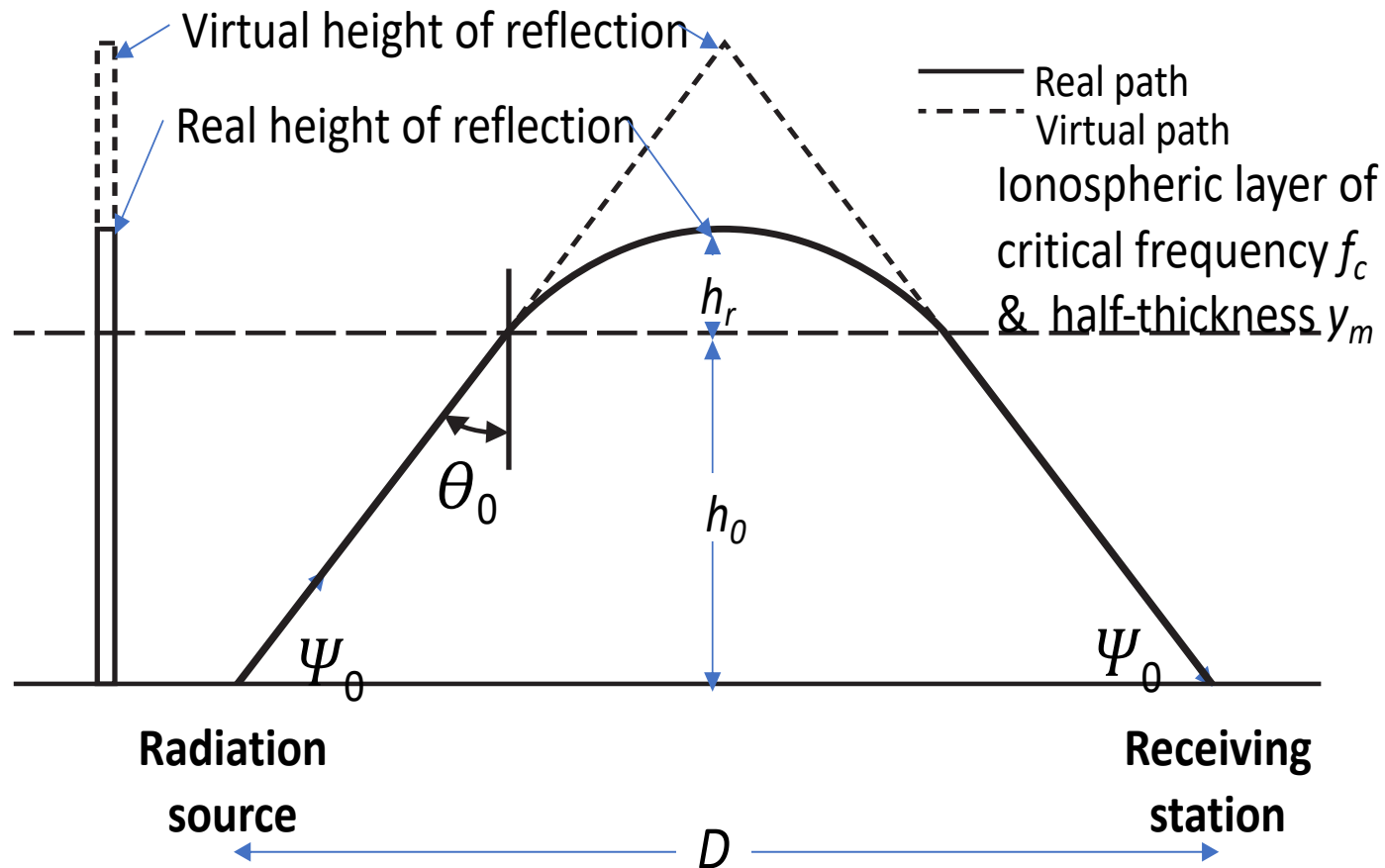
- Zoomed-in views of TeePees often appear diffuse with short discrete bursts over limited frequency ranges.
- Diffuse appearance may imply significant ionospheric scattering associated with moderate geomagnetic activities ($K_p \sim 4$).



<https://omniweb.gsfc.nasa.gov/>

http://wdc.kugi.kyoto-u.ac.jp/ae_realtime/201910/index_20191024.html

Hypothesis: Lightning as TeePee Source



Breit & Tuve Theorem [1926]

- Successive refraction can be regarded as a single reflection from a virtual reflection point at a virtual height.

Ground Range as a Function of Wave Frequency (f/f_c) and Incident Angle (θ_0)

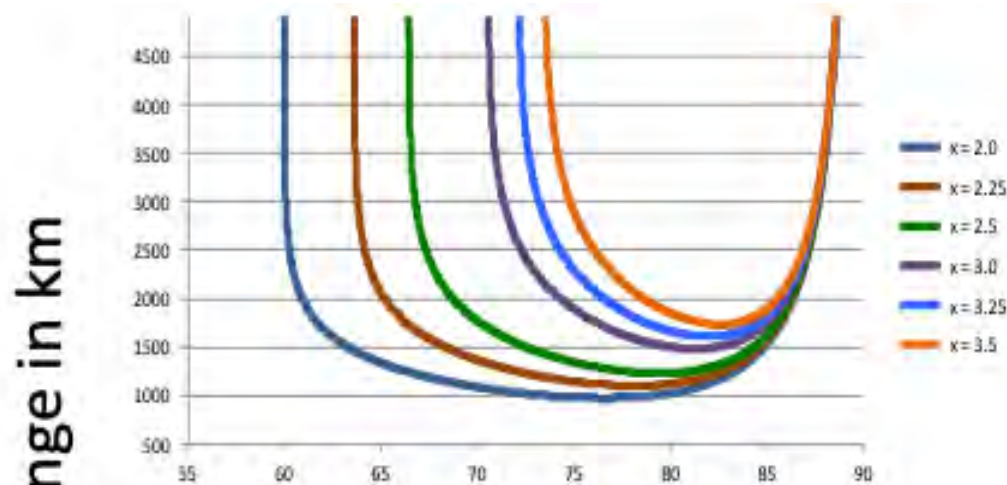
- For flat earth surface and horizontally stratified ionosphere

$$D = 2h_0 \tan \theta_0 + \frac{f}{f_c} y_m \sin \theta_0 \ln \left[\frac{\left(1 + \frac{f}{f_c} \cos \theta_0\right)}{\left(1 - \frac{f}{f_c} \cos \theta_0\right)} \right] \quad y_m = \left\{ \left[\left(f_p^{-1}\right) \partial^2 f_p(h) / \partial h^2 \right]_{h_c} \right\}^{-1/2} \approx h_c - h_0$$

- For spherical earth surface and concentric ionospheric layers

$$D \approx 2R_E \left\{ \frac{\pi}{2} - \theta_0 - \cos^{-1} \left[\frac{(R_E + h_0) \sin \theta_0}{R_E} \right] \right\} + \frac{x y_m R_E \sin \theta_0}{(R_E + h_0)} \ln \left[\frac{(1 - \alpha) + \gamma}{(1 - \alpha) - \gamma} \right]$$

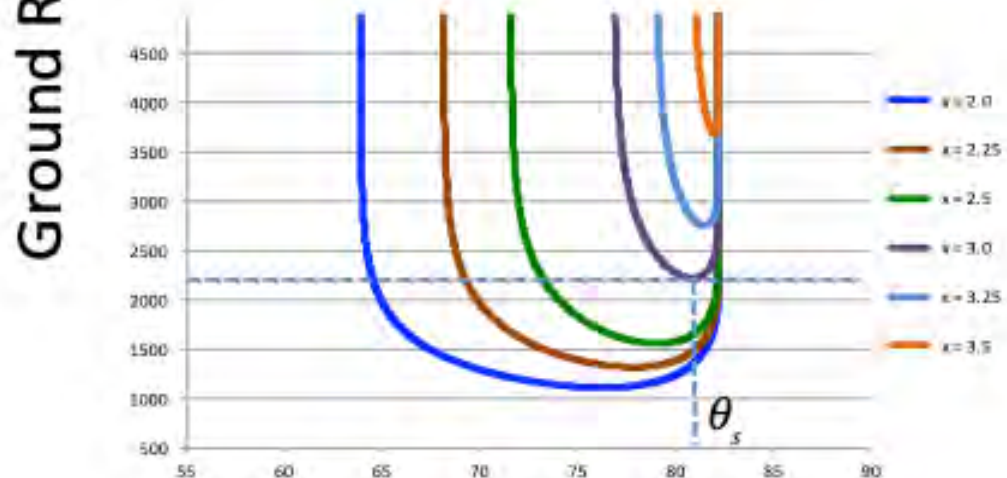
Ground Range as a Function of Wave Frequency (f/f_c) and Incident Angle (θ_0)



Flat Earth

For any U-curve at a given frequency, $x = f/f_c$

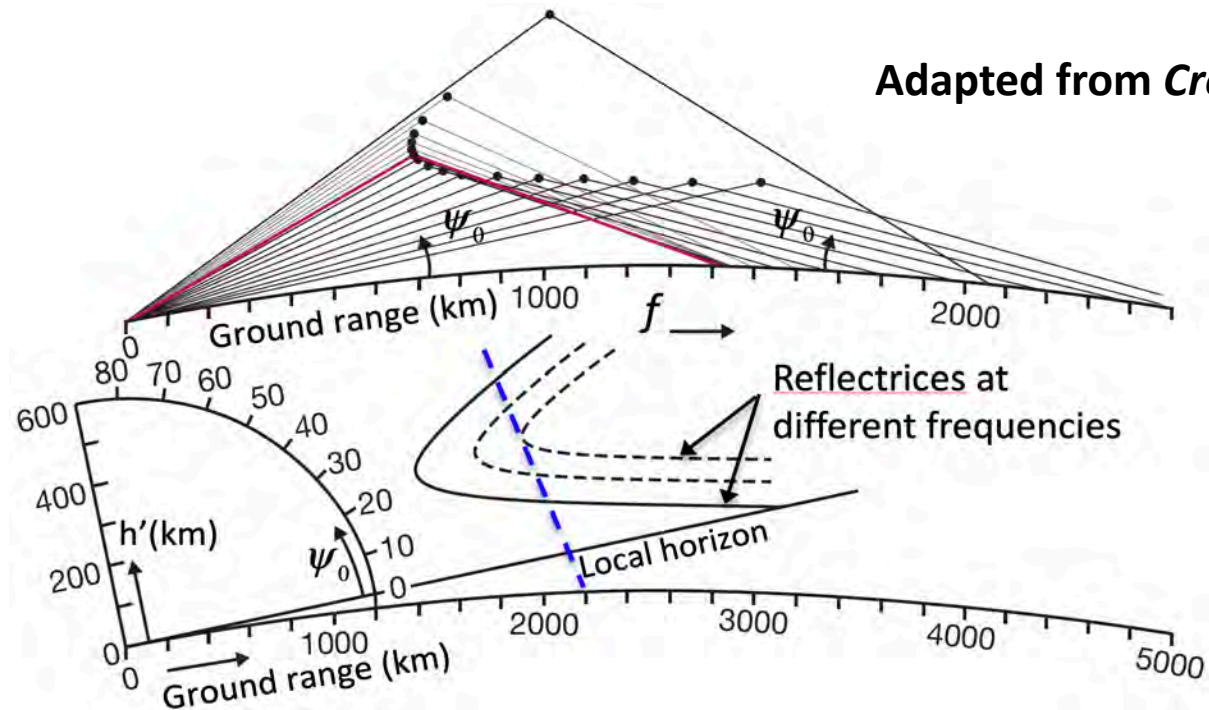
- The U-minimum gives the MUF(D) and MUF ray direction, θ_s .
- Rays with $\theta_0 > \theta_s$ are called low-angle rays. One with $\theta_0 < \theta_s$ are high-angle rays.



Spherical Earth

Incidence Angle wrt the Vertical @ h_0

2-D Ray Tracing Calculations



- For a fixed frequency f , the loci of the virtual reflection point as a function of ψ_0 ($= 90^\circ - \theta_0$) is the **reflectrix**.
- Each reflectrix corresponds (per frequency) to a U-curve.
- TeePee spectral signature results from thunderstorm motion (*in any direction*) relative to the fixed observing location.

Summary

- TeePee signature can be explained in terms of ionospheric reflection of HF emissions from remote lightning storms.
- Nested TeePees may be due to line of lightning storms moving *along* antenna beam.
- TeePee series may be due to line of thunderstorms moving *across* antenna beam.
- TeePee observations can provide a means to study remote ionospheric conditions (h_o , y_m , f_c).
- TeePee fine structures can help deduce scale sizes of bottom-side ionospheric irregularities.

