

HamSCI and the 2017 Total Solar Eclipse

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Outline

- I. What is Ham Radio & HamSCI?**
- II. Eclipse Experiments**
 - I. 2017 Total Solar Eclipse &
The Ionosphere**
 - II. Ham Radio Data Sources**
 - III. Observations & Discussion**
- III. Results & Summary**



Amateur/Ham Radio

- Hobby for Radio Enthusiasts
 - Communicators • Builders • Experimenters
- Wide-reaching Demographic
 - All ages & walks of life
 - Over 730,000 US hams; ~3 million World Wide

[<http://www.arrl.org/arrl-fact-sheet>]



HamSCI
<http://hamsci.org>

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HamSCI

The Ham radio Science Citizen Investigation is:



hamsci.org/dayton2017

An organization that allows university researchers to collaborate with the amateur radio community in scientific investigations.

Objectives:

1. **Advance** scientific research and understanding through amateur radio activities.
2. **Encourage** the development of new technologies to support this research.
3. **Provide** educational opportunities for the amateur community and the general public.



Founder/Lead HamSCI Organizer:
Dr. Nathaniel A. Frissell, W2NAF
NJIT Center for Solar-Terrestrial Research



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Total Solar Eclipse

21 August 2017

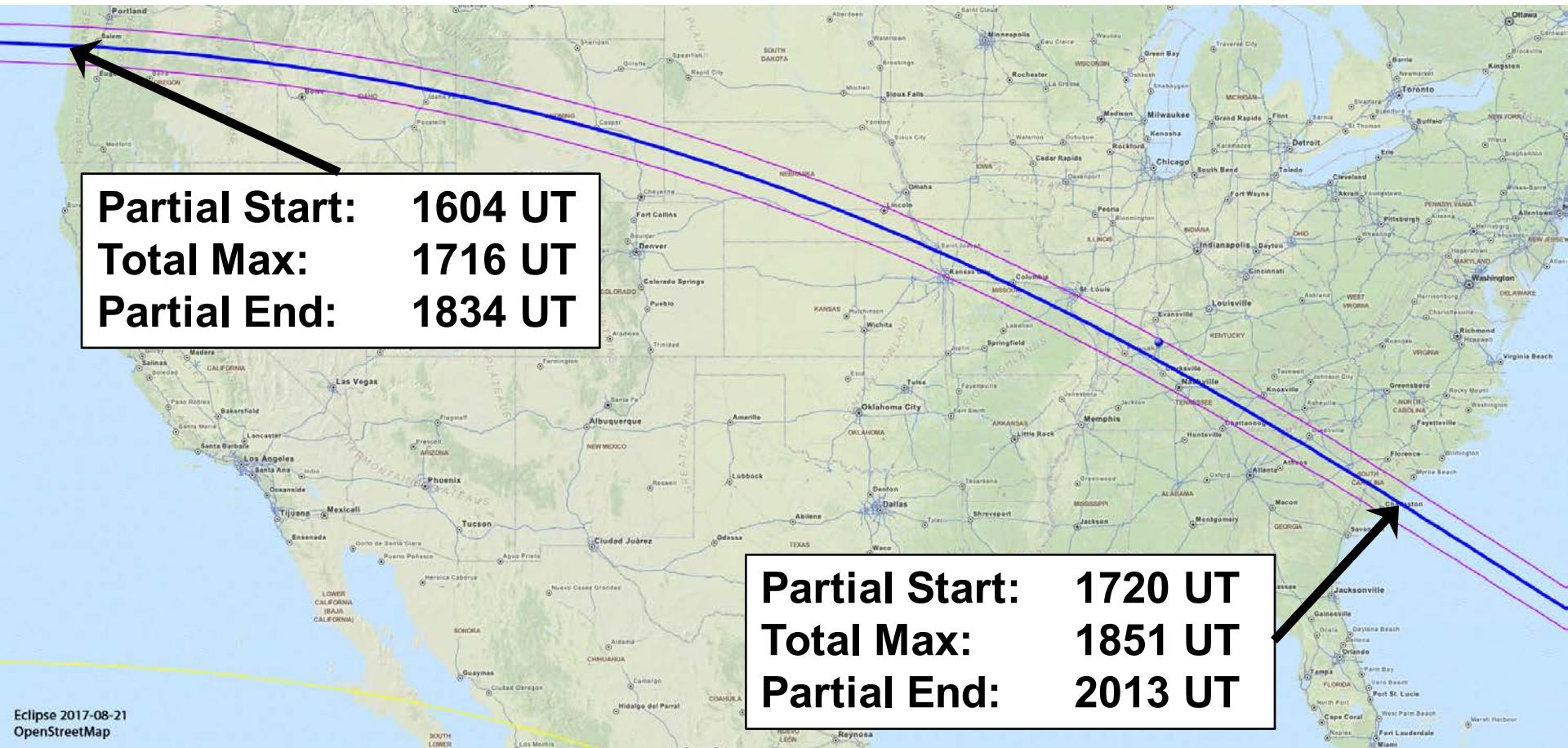
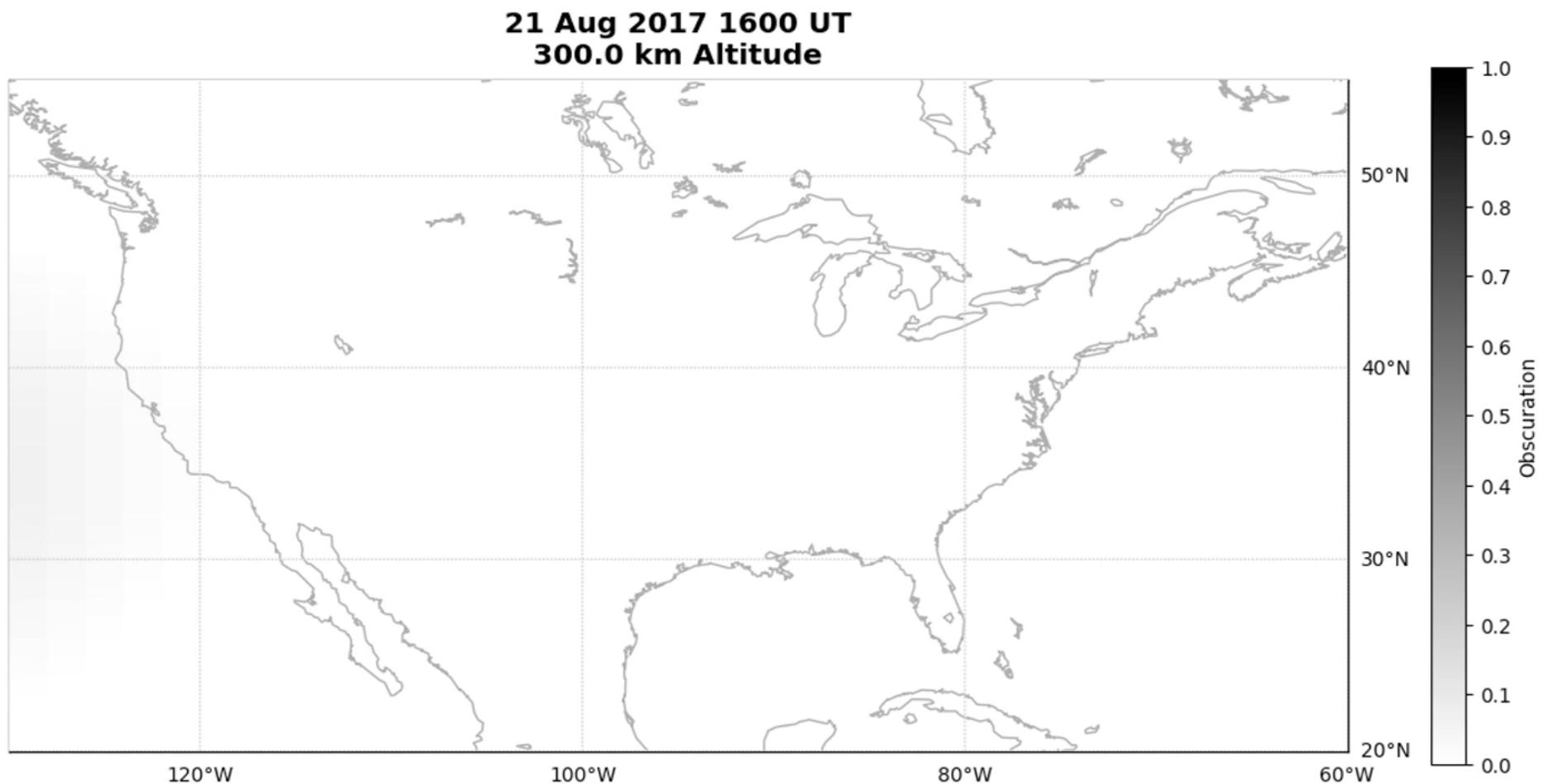


Figure: W. Strickling, Wikipedia



Eclipse Obscuration



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Solar Eclipse Ionospheric Effects?

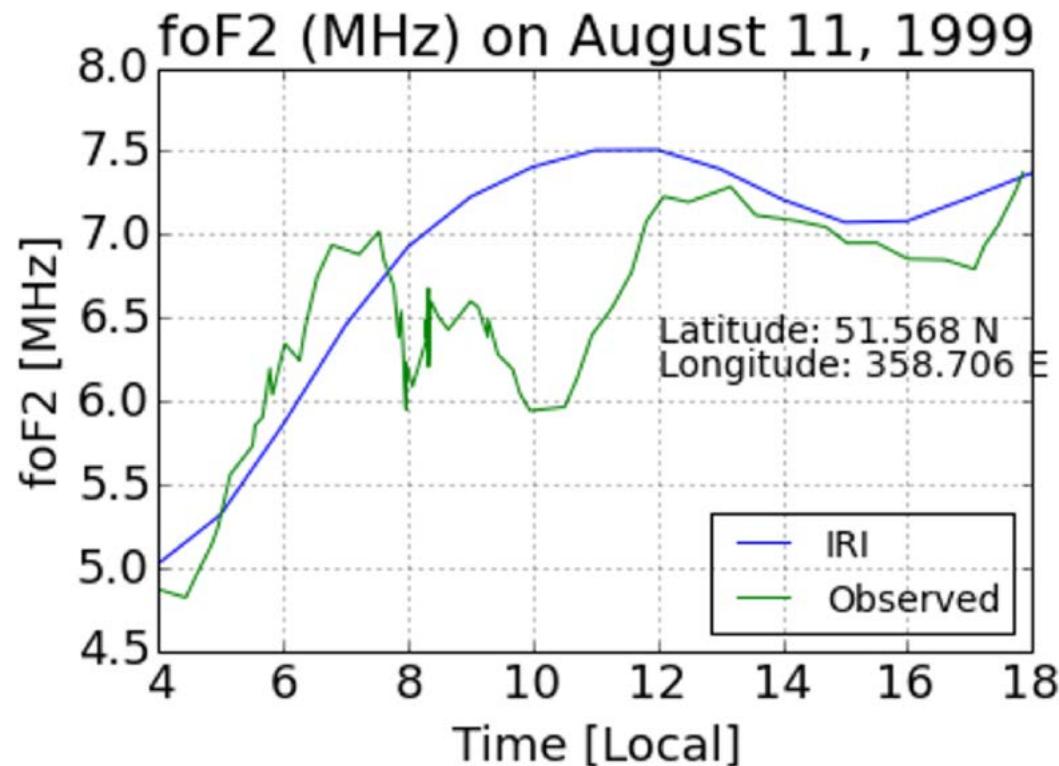
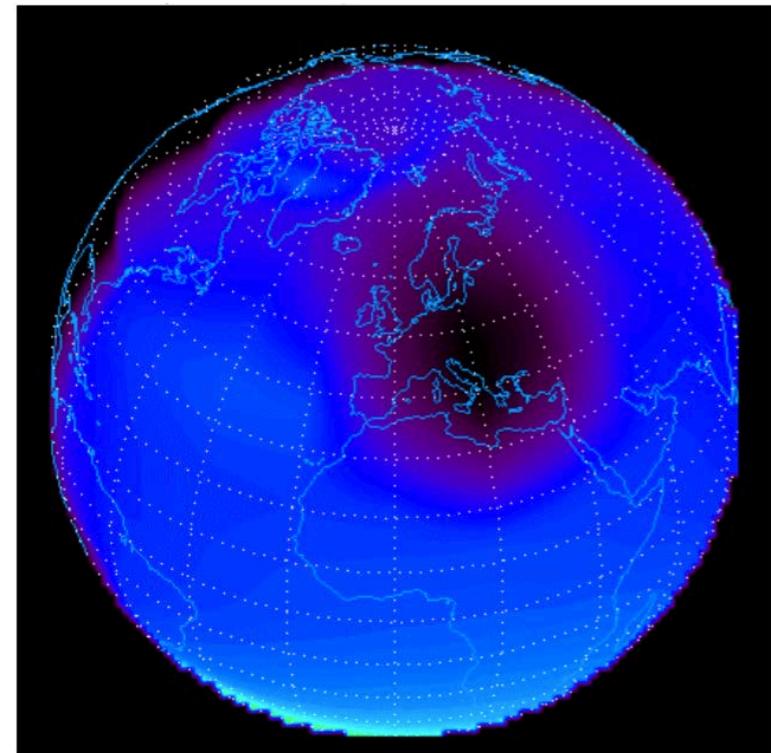


Figure: M. Moses after
Afraimovich *et al.*, 2002



Model Electron Density at ~280
km alt. during 1999 Eclipse
M. Harris from Bamford 2000.



HamSCI Eclipse Research Questions

- What are the temporal and spatial scales of eclipse-induced ionospheric effects?
- Can we observe TIDs in the ionosphere caused by the eclipse?
- How does the eclipse affect HF propagation?



Amateur Radio and the HF Bands

Frequency	Wavelength
1.8 MHz	160 m
3.5 MHz	80 m
7 MHz	40 m
10 MHz	30 m
14 MHz	20 m
18 MHz	17 m
21 MHz	15 m
24 MHz	12 m
28 MHz	10 m
50 MHz	6 m



K2MFF, The NJIT Ham Radio Station

- Hobbyists routinely use HF-VHF transionospheric links.
- Often ~100 W into dipole antennas.

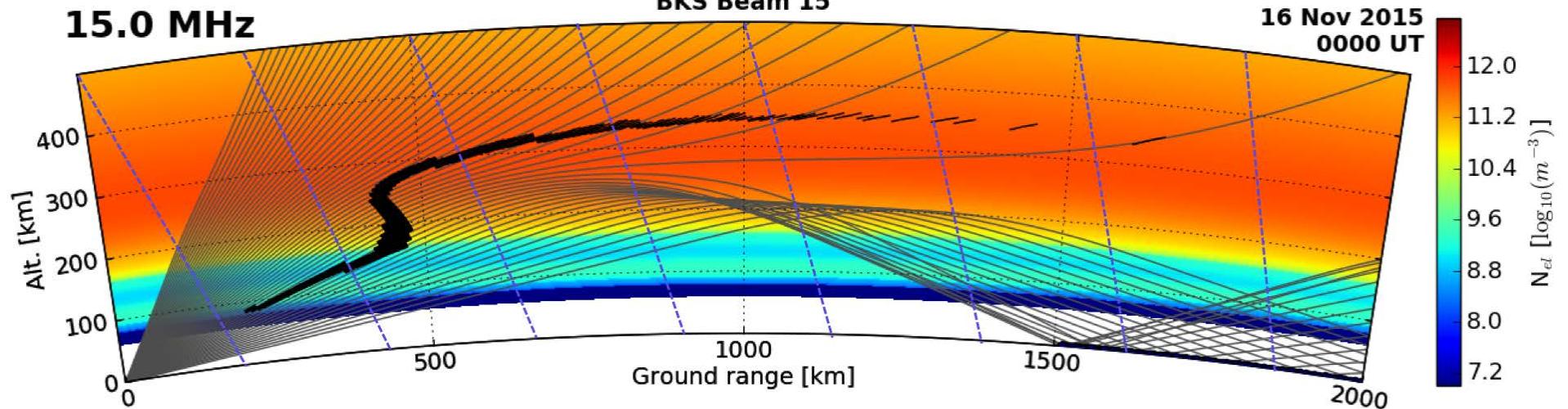


HF Propagation & The Ionosphere

15.0 MHz

BKS Beam 15

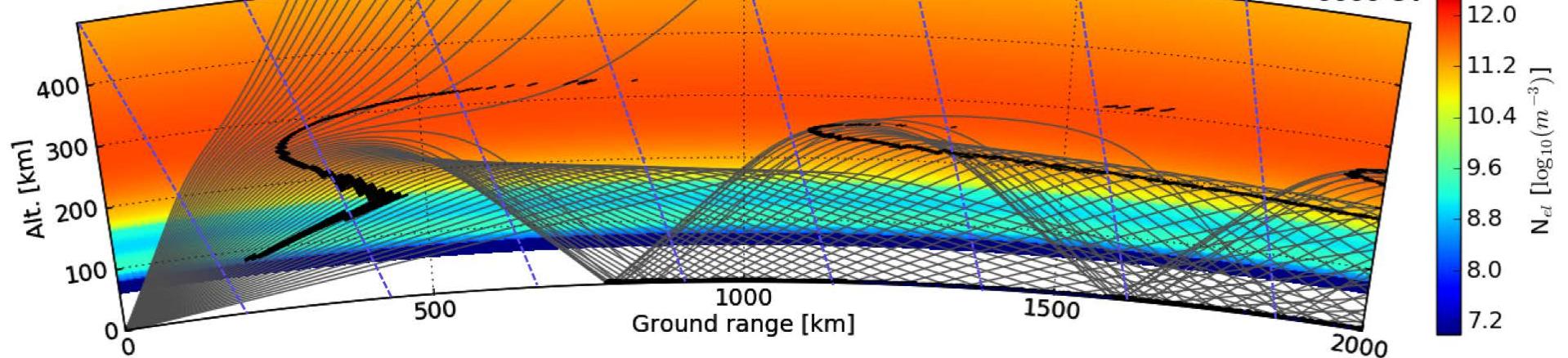
**16 Nov 2015
0000 UT**



10.0 MHz

BKS Beam 15

**16 Nov 2015
0000 UT**



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HamSCI Eclipse Experiments

- **Solar Eclipse QSO Party (SEQP)**

- Ham Radio Contest-Like Event
- Generate a quasi-random dataset
- Data from RBN, PSKRepoter, WSPRNet, Logs

- **HF Wideband Recording**

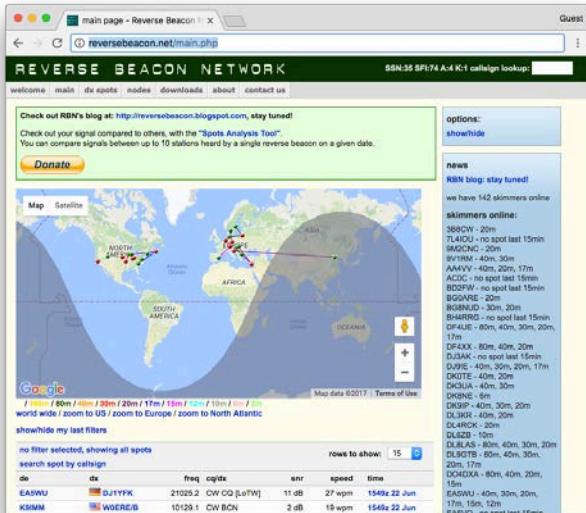
- Use SDRs to record large amounts of HF Spectrum

- **HF Frequency Measurement Experiment**

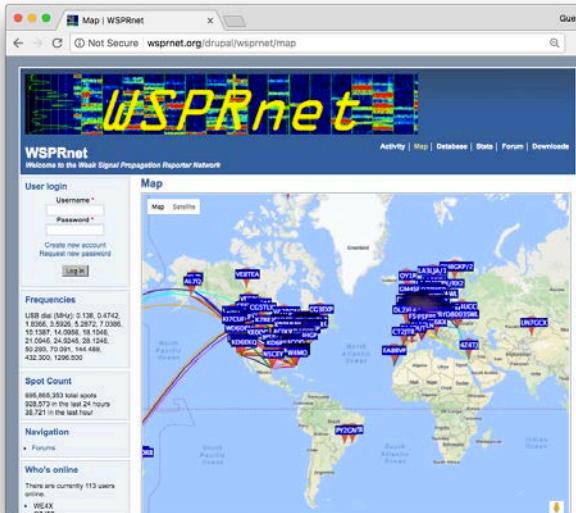
- Measure changes in WWV, CHU frequency due to eclipse



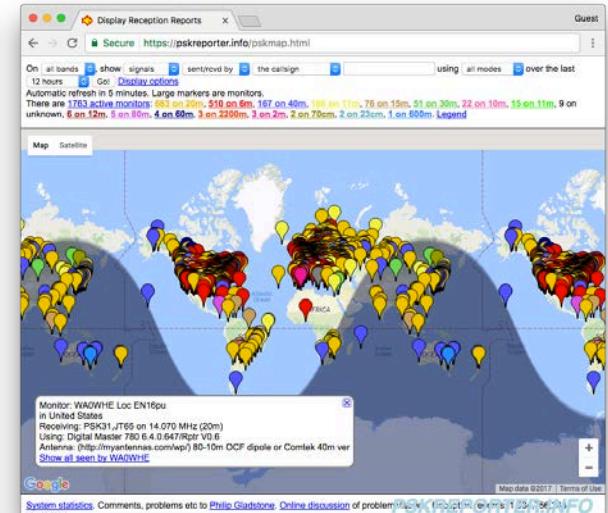
Big Data – Ham Radio Observatories



RBN
reversebeacon.net



WSPRNet
wsprnet.org



PSKReporter
pskreporter.info

Network	Start Year	# Spots	DB Size
RBN	2009	578,000,000	36 GB
WSPRNet	2008	535,000,000	44 GB
PSKReporter	2013	1,000,000,000	100 GB

There is A LOT of data!



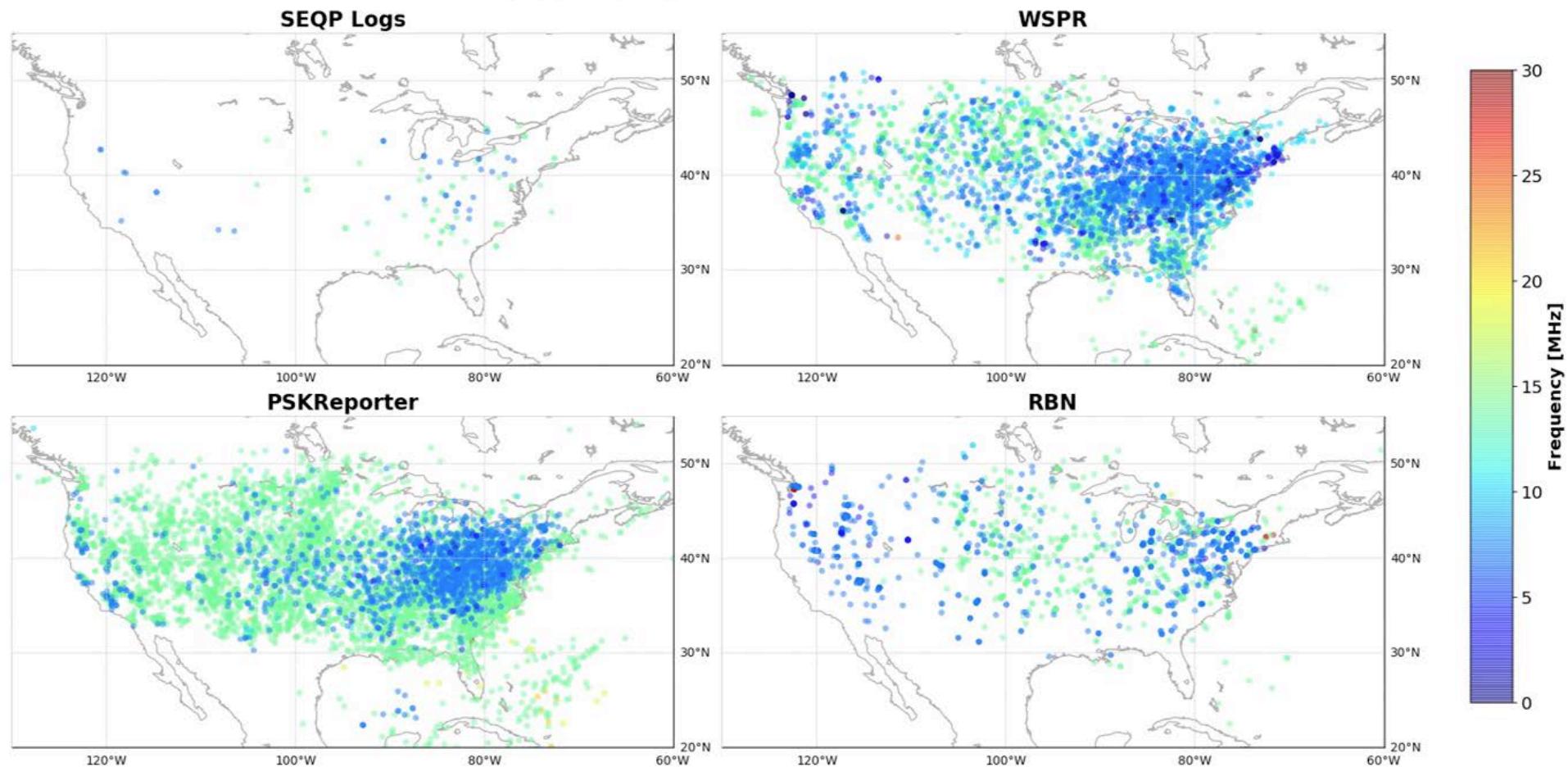
Data Collection

hamsci.org	<ul style="list-style-type: none">• 600+ Parsed Logs• 30,768 QSOs
zenodo.org HamSCI Community	<ul style="list-style-type: none">• 50+ Submissions
Reverse Beacon Network	<ul style="list-style-type: none">• 618,623 Spots
PSKReporter	<ul style="list-style-type: none">• 1,287,855 Spots
WSPRNet	<ul style="list-style-type: none">• 630,132 Spots



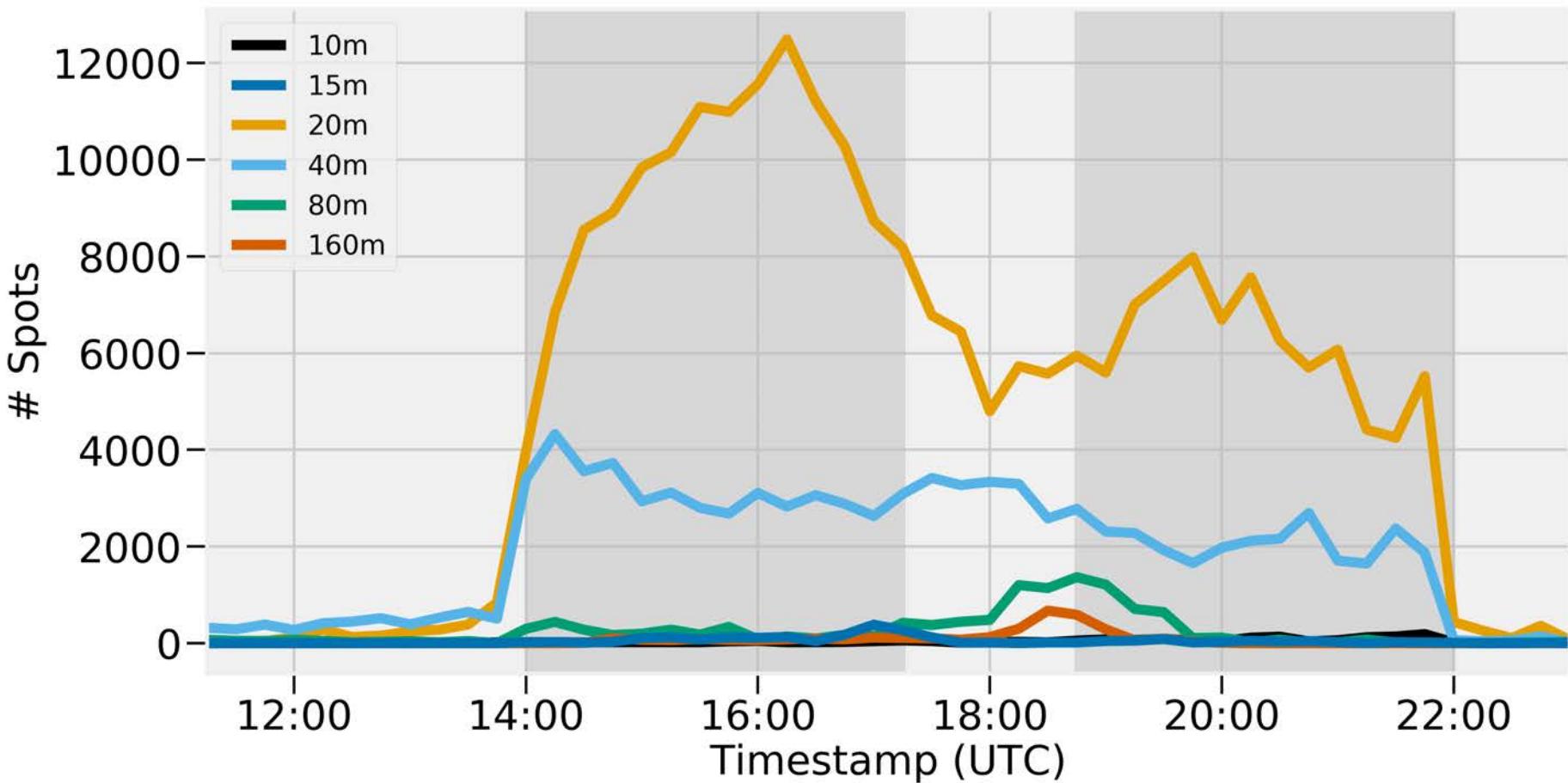
Ham Radio Eclipse Data

21 Aug 2017 1400 UT - 21 Aug 2017 1405 UT
QSO/Spot Midpoints; 300.0 km Obscuration Alt



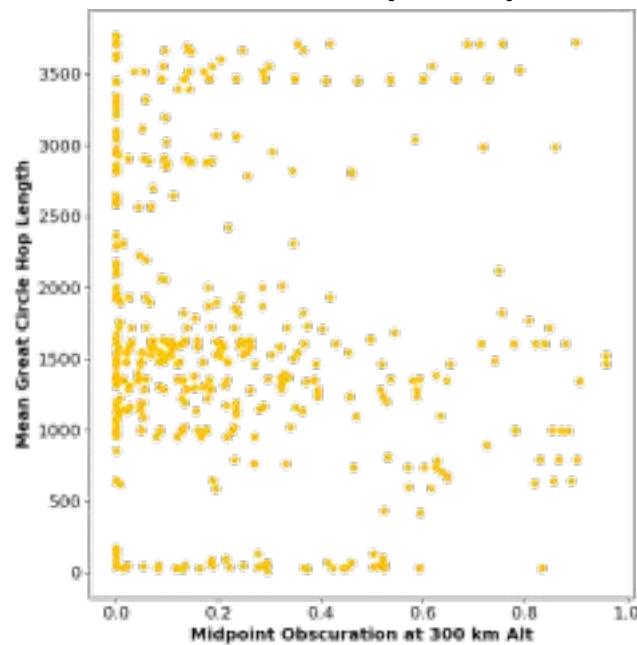
SEQP RBN Spots

RBN SEQP Spots by Band (Contiguous US TX and RX Only)

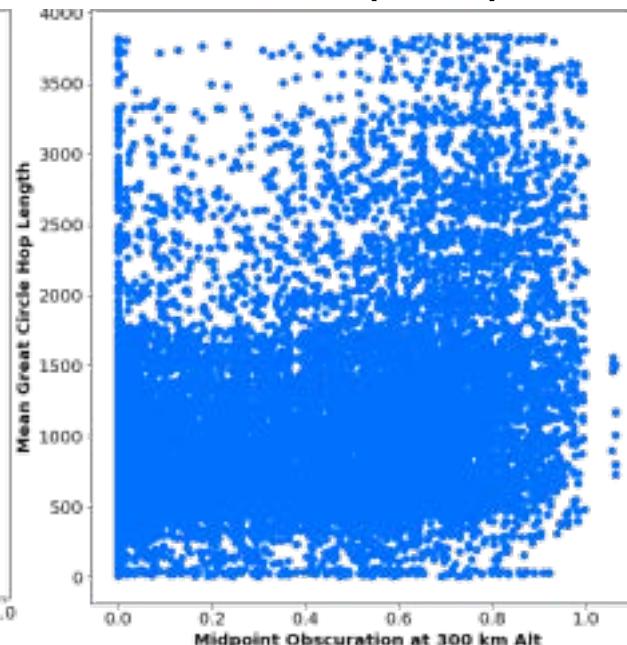


Hop Length vs. Obscuration

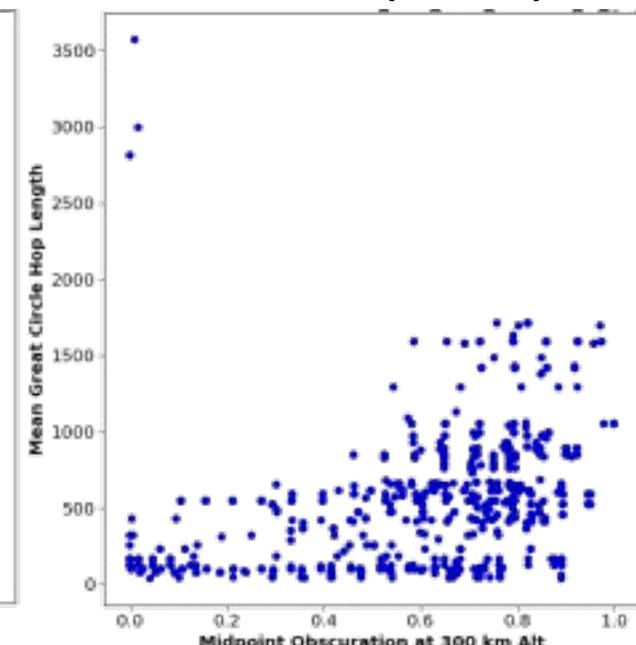
21 MHz (15 m)



7 MHz (40 m)



1.8 MHz (160 m)



Greater Eclipse → Fewer Spots

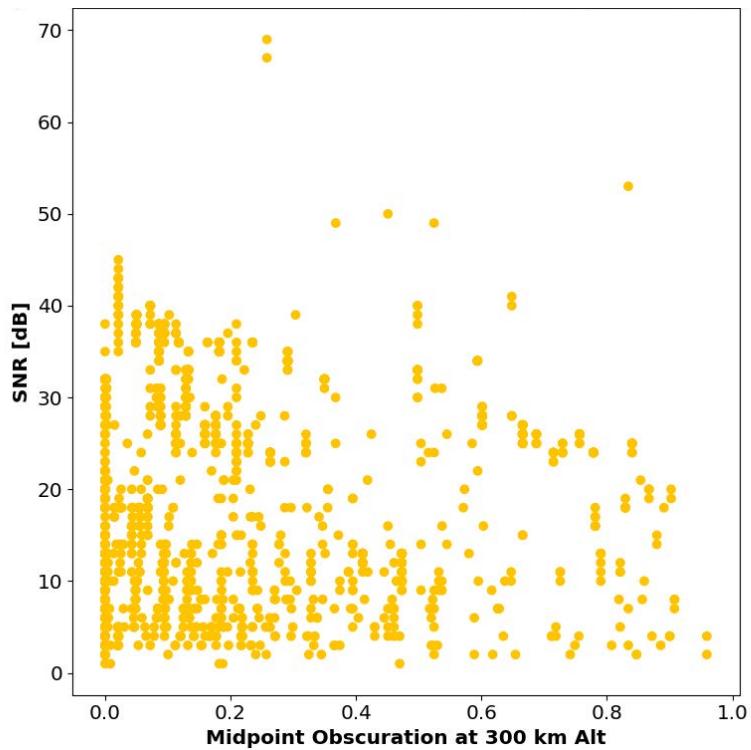
Greater Eclipse → Longer Hops

Greater Eclipse → More & Longer!

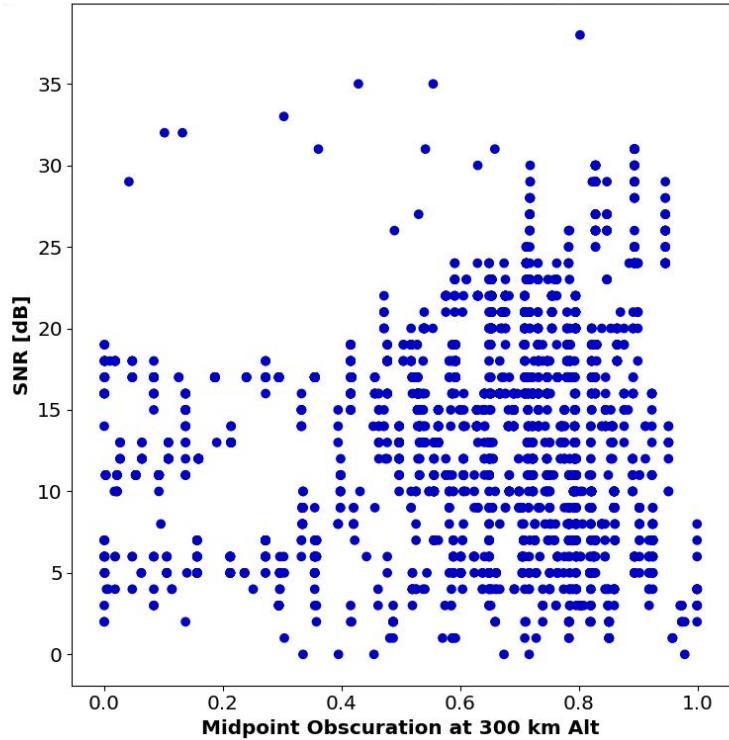


SNR vs. Obscuration

21 MHz (15 m)



1.8 MHz (160 m)



Weaker F Region →
Lower SNR, Fewer Spots

Weaker D Region →
Less Absorption, Stronger Signals,
& More Spots



Summary

- **Ham Radio Science Citizen Investigation**

- An organization that allows university researchers to collaborate with the amateur radio community in scientific investigations.

- **2017 Total Solar Eclipse**

- Shadow of eclipse stops ion production in ionosphere
- During Eclipse:
 - Number of 1.8 to 7 MHz Spots Increased
 - Number of 14 to 30 MHz Spots Decrease
 - Hop Lengths & SNRs increased on lower bands < 14 MHz
 - SNRs decreased on higher bands (> 14 MHz)
- SEQP observations suggest raising of the F layer and depletion of the D layer.



Thank you!



References

- Afraimovich, E.L., E.A. Kosogorov, O.S. Lesyuta (2002), Effects of the August 11, 1999 total solar eclipse as deduced from total electron content measurements at the GPS network, Journal of Atmospheric and Solar-Terrestrial Physics, Volume 64, Issue 18, Pages 1933-1941, ISSN 1364-6826, [http://dx.doi.org/10.1016/S1364-6826\(02\)00221-3](http://dx.doi.org/10.1016/S1364-6826(02)00221-3).**
- Bamford, R. (2000), Radio and the 1999 UK Total Solar Eclipse, Rutherford Appleton Laboratory, Chilton, Didcot, UK.**

